Metastable Vacua from SCFTs Based on arXiv:1003.0523 with A. Amariti, A. Mariotti and M.Siani

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Outline

SQCD and Seiberg duality The IR free magnetic window: ISS model ISS model in the Conformal window Metastable vacua in the Superconformal SSQCD Generalizations Conclusions

Outline

SQCD and Seiberg duality

The IR free magnetic window: ISS model

ISS model in the Conformal window

Metastable vacua in the Superconformal SSQCD

Generalizations

Conclusions

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► $\mathcal{N} = 1$ SUSY $SU(N_c)$ gauge theory, $N_f Q$ and \tilde{Q} with $SU(N_f)^2$ sym. $\mathcal{L}_{SQCD} = \int d^4\theta K_{ij} \left(Q_i e^{-V} Q_i^{\dagger} + \tilde{Q}_i e^{V} \tilde{Q}_i^{\dagger} \right) + \frac{1}{2a^2} \int d^2\theta W_{\alpha} W^{\alpha} + \int d^4\theta W(Q, \tilde{Q}) + \text{h.c.}$



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Seiberg duality

If $N_f > N_c + 1$ the *electric* SQCD theory in the IR is equivalently described by a *magnetic* theory

	Electric theory	Magnetic Theory				
Gauge	$SU(N_c)$	$SU(N_f - N_c)$				
Flavor	$SU(N_f)$	$SU(N_f)$				
Fields	$oldsymbol{Q}, ilde{oldsymbol{Q}}$	$q, ilde{q}$ and $M = rac{Q ilde{Q}}{\Lambda}$				
Superpotential	W = 0	$W = Mq\tilde{q}$				
Mass terms	$\Delta W = mQ ilde{Q}$	$\Delta W = \mu^2 M$				
Holomorphic scale	$\Lambda_h = Ee^{rac{-8\pi^2}{bg^2}}$	$ ilde{\Lambda}_{h}={\it Ee}^{rac{-8\pi^{2}}{ ilde{b} ilde{g}^{2}}}$				
$\underline{\text{Scale matching}} \ \text{(mapping of the SUSY vacua)} \qquad \Lambda_h^b \tilde{\Lambda}_h^{\tilde{b}} = (-1)^{2N_f} \hat{\Lambda}^{2N_f}$						
with $b=3N_c-N_f$, $\tilde{b}=2N_f-3N_c$ and $\hat{\Lambda}$ is an intermediate scale						

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UV/IR duality maps a strongly coupled theory to a weakly coupled one

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$$\blacktriangleright$$
 $N_c+1 < N_f < rac{3}{2}N_c$ $b>0$ and $ilde{b} < 0$

UV free electric window - IR free magnetic window

$$\boxed{\frac{3}{2}N_c < N_f < 3N_c}$$
 Conformal Window

IR interacting fixed point in both sides of duality

 $N_f > 3N_c$

$$\begin{array}{l} \text{If } N_f < 2N_c \rightarrow \tilde{g} < g \\ \text{If } N_f > 2N_c \rightarrow \tilde{g} > g \\ < 0 \text{ and } \tilde{b} > 0 \end{array}$$

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IR free electric window - UV free magnetic window

- SUSY breaking in the IR free magnetic window
- $\blacktriangleright W_{ele} = mQ\tilde{Q}, \qquad N_c + 1 < N_f < \frac{3}{2}N_c$
- Witten Index: N_c SUSY vacua
- Strongly coupled IR physics \rightarrow not accessible
- UV/IR duality, magnetic description

$$W = h M q \tilde{q} - \mu^2 M$$

where

$$\mu \sim -m\hat{\Lambda}, \quad h \sim rac{\Lambda}{\hat{\Lambda}}, \quad M \sim rac{Q\tilde{Q}}{\Lambda}$$

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Kahler canonical, correction suppressed

Tree level SUSY breaking

• Solving equation of motion: $F_M = h q^i_\alpha \tilde{q}^\alpha_j - h \mu^2 \delta^i_j$

with
$$\alpha = 1, \ldots, N_f - N_c$$
, $i, j = 1, \ldots, N_f$

- ► Rank Condition: $F_M = h\mu^2 \delta_{j'}^{i'}$ with $i', j' = N_f N_c + 1, \dots, N_f$
- At tree level SUSY is broken, and $V_{tree} = N_c |h\mu^2|^2$

► (Pseudo)-moduli space $q = \begin{pmatrix} \mu e^{\theta} \\ 0 \end{pmatrix}, \quad \tilde{q}^{T} = \begin{pmatrix} \mu e^{-\theta} \\ 0 \end{pmatrix}, \quad M = \begin{pmatrix} 0 & 0 \\ 0 & X \end{pmatrix}$

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- Tree level analysis inconclusive, one loop analysis necessary
- ▶ Ignoring corrections from the Kahler $(\gamma_i \sim \frac{h^2}{32\pi^2})$
- Effective potential: CW

$$V_{eff}^{(1)} = \frac{1}{64\pi^2} STr\mathcal{M}^4 \log \frac{\mathcal{M}^2}{\Lambda^2} = \frac{1}{64\pi^2} STr\left(Trm_B^4 \log \frac{m_B^2}{\Lambda^2} - Trm_F^4 \log \frac{m_F^2}{\Lambda^2}\right)$$

Fluctuations around the vacuum

$$q = \begin{pmatrix} \mu e^{\theta} + \sigma_1 \\ \phi_1 \end{pmatrix}, \quad \tilde{q}^{\mathsf{T}} = \begin{pmatrix} \mu e^{-\theta} + \sigma_2 \\ \phi_2 \end{pmatrix}, \quad M = \begin{pmatrix} \sigma_3 & \phi_3 \\ \phi_4 & X \end{pmatrix}$$

Relevant superpotential

$$W = X\phi_1\phi_2 - h\mu^2 X + h\mu e^{-\theta}\phi_1\phi_4 + h\mu e^{\theta}\phi_2\phi_3$$

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Mass matrices

$$m_B^2 = \begin{pmatrix} W^{\dagger ac} W_{cb} & W^{\dagger abc} W_c \\ W_{abc} W^{\dagger c} & W_{ac} W^{\dagger cb} \end{pmatrix} \qquad m_f^2 = \begin{pmatrix} W^{\dagger ac} W_{cb} & 0 \\ 0 & W_{ac} W^{\dagger cb} \end{pmatrix}$$

- Classical spectrum with $m_f^2 = m_f^2(X)$ and $m_b^2 = m_b^2(X)$
- Minimum of the CW effective potential at

$$\langle X
angle = 0$$
 & $\langle heta + heta^*
angle = 0$

$$m_X^2 = N_c \frac{h^2 \mu^2}{8\pi^2} (\log 4 - 1)$$

$$m_{ heta+ heta^*}^2 = (N_f - N_c) rac{n^- \mu^-}{8\pi^2} (\log 4 - 1)$$

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- Electric theory: N_c SUSY vacua
- Duality $\rightarrow N_c$ SUSY vacua also in the magnetic model
- Indeed at large $\langle M \rangle$ the q and \tilde{q} quarks are integrated out
- Scale matching $\Lambda_h^{\tilde{b}} = \Lambda_L^{\tilde{b}-N_f}$ det hM
- ▶ The effective theory at large *M* is pure SYM with singlets

$$W = W_{dyn} + W_{tree} = (N_f - N_c)\Lambda_L^3 - h\mu^2 M$$

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with

$$W_{dyn} = (N_f - N_c) \left(ilde{\Lambda}_h^{ ilde{b}} \det(hM)
ight)^{rac{1}{N_c}}$$

$$\blacktriangleright \ \langle hM \rangle_{SUSY} = \mu \left(\frac{\mu}{\tilde{\Lambda}_h} \right)^{\frac{b}{N_c}} \gg \mu \qquad \text{if} \qquad \tilde{\Lambda}_h \gg \mu$$

• Lifetime:
$$\Gamma \sim e^{-S_B}$$

For a scalar field the classical bounce action is

$$S_B = \int d^4x \, \partial_\mu \phi \, \partial^\mu \phi + V(\phi)$$

- If $V(\phi)$ is a triangular barrier $S_B \sim rac{(\Delta \phi)^4}{\Delta V}$
- $\blacktriangleright \ \text{Here} \ S_B \sim \left(\frac{\mu}{\tilde{\Lambda}_h} \right)^{\frac{4\tilde{b}}{N_c}} \gg 1 \quad \text{ with } \quad \tilde{b} < 0 \ \text{and} \ \tilde{\Lambda}_h \gg \mu$
- ➤ → Parametrically large lifetime (the more the theory is weakly coupled the more the lifetime is large)

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Outline SQCD and Seiberg duality The IR free magnetic window: ISS model ISS model in the Conformal sSQCD Generalizations Conclusions

- Massive SQCD with $\frac{3}{2}N_c < N_f < 3N_c$
- Electric and magnetic theories have are interacting SCFTs
- $\tilde{g} < g$ if $N_f < 2N_c$
- Classically supersymmetry is broken as in the IR free window

• But naively
$$S_B = \left(\frac{\mu}{\tilde{\Lambda}_b}\right)^{\frac{b}{N_c}}$$
 is not large

$$\tilde{b} = 3(N_f - N_c) - N_f = 2N_f - 3N_c > 0$$

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• Equivalently: W_{dyn} is not negligible at small M

Deeper analysis

There are two important differences with the IR free case

- The $\gamma_i \neq 0 \rightarrow Z_i \neq 1$: RG effects not negligible
- ▶ RG flow from Λ_{UV} to $\Lambda_{IR} \equiv \Lambda_c$, CFT exit scale

▶ When $\gamma_i \sim cost. \neq 0$: holomorphic → canonical basis

$$\begin{array}{ll} \underline{\mbox{Holomorphic basis}} & \underline{\mbox{Canonical basis}} \\ \hline \phi_i & \phi_i' = \sqrt{Z}_i \phi_i \\ K = Z \phi_i \phi_i^{\dagger} & K = \phi_i' \phi_i'^{\dagger} \\ W = W(\lambda_{\textit{bare}}, \phi_i) & W = W(\lambda_{\textit{phys.}}, \phi_i') \end{array}$$

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- If $Z_{UV} = 1$ in the UV, then $Z_{IR} \neq 1$ after RG flow.
- Canonical basis: absorb the WFR in the fields at every scale.
- RG effects on the couplings in W: from

$$\Delta W = \lambda_i(E)\phi_i$$

with $\lambda(E) = \Lambda(E')Z(E, E')^{-\frac{\gamma_i}{2}}$

Running of the holomorphic scale

$$\Lambda_h(E') = \Lambda_h(E) \prod z_{\phi_i}^{T_{F_i}/2}$$

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In the canonical basis $\Lambda_h(E') = \Lambda_h(E) \frac{E'}{E}$

• Renormalization of S_B : $S_B(E') = S_B(E)Z_M^3$

► CFT exit scale: two possible sources, $\langle M \rangle_{susy}$ or μ ► From F_M $M = \mu(\Lambda_{IR}) \left(\frac{\tilde{\Lambda}_h(\Lambda_{IR})}{\mu(\Lambda_{IR})}\right)^{\frac{\tilde{b}}{N_c}}$

• If
$$M = \Lambda_c$$
, $\tilde{\Lambda}_h = \Lambda_c e^{-\frac{8\pi^2}{g^2\tilde{b}}}$ and

$$\mu(\Lambda_c) = e^{-\frac{4\pi^2}{g^{*^2(N_f - N_c)}}} \Lambda_c \ll \Lambda_c$$

(If $\mu(\Lambda_C) = \Lambda_c$ then $M \gg \Lambda_c$)

$$S_B = \left(\frac{\mu_{IR}}{\tilde{\Lambda}_{h,IR}}\right)^{4\frac{2N_f - 3N_c}{N_f - N_c}} = e^{\frac{16\pi^2}{g_*^2(N_f - N_c)}}$$

The lifetime cannot be parametrically large in the ISS model in the CFT window. Metastable vacua cannot be long lifetime.

- Is it possible to overcome this problem?
- Yes, but other relevant deformations must be added
- Indeed in ISS there are only two scales
 - μ the SUSY breaking scale
 - *M* the SUSY restoration scale

and they are proportional.

- S_B is their ratio, and cannot be made parametrically large
- ▶ With other deformations S_B can be parametrically large OR equivalently

The non perturbative potential can be negligible at small vevs.

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Example: the SSQCD

▶ $\mathcal{N} = 1$ $SU(N_c)$ gauge theory with $SU(N_1) \times SU(N_2)$ flavour symmetry, N_1 massive and N_2 massless fundamentals, and a bifundamental of $SU(N_2)$

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	$SU(N_c)$	$SU(N_1)$	$SU(N_2)$
$oldsymbol{Q}\oplus ilde{oldsymbol{Q}}$	$N_c\oplusar{N}_c$	$ar{N}_1 \oplus N_1$	1
$P\oplus ilde{P}$	$N_c\oplusar{N}_c$	1	$ar{N}_2\oplus N_2$
S	1	1	$N_2\otimes ar{N}_2$
Superpote	ential: W =	$= mQ\tilde{Q} + S$	$SP\tilde{P} + k^2S$

Seiberg duality: if $N_1 + N_2 > N_c + 1$ there is a dual $SU\left(\widetilde{N} = N_1 + N_2 - N_c\right)$ dual gauge theory $\begin{array}{c|c} & SU(\widetilde{N}) & SU(N_1) & SU(N_2) \\ \hline q \oplus \widetilde{q} & \widetilde{N} \oplus \overline{\widetilde{N}} & \overline{N}_1 \oplus N_1 & 1 \\ \hline p \oplus \widetilde{p} & \widetilde{N} \oplus \overline{\widetilde{N}} & 1 & \overline{N}_2 \oplus N_2 \\ M & 1 & N_1 \otimes \overline{N}_1 & 1 \\ L \oplus K & 1 & N_1 \oplus \overline{N}_1 & \overline{N}_2 \oplus N_2 \end{array}$ $\begin{array}{c|c} W = Mq\widetilde{q} + Lq\widetilde{p} + Kp\widetilde{q} - \mu^2 M + \rho p\widetilde{p} \end{array}$

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- Conformal window: $3/2N_c < N_1 + N_2 < 3N_c$
- Another constraint imposed by the rank condition: Tree level SUSY breaking if $N_1 > \widetilde{N} \rightarrow N_2 < N_c$
- ▶ Defining $n_{ele} = \frac{N_2}{N_1}$ and $x_{ele} = \frac{N_c}{N_1}$ there is no meson hitting the unitary bound $(R(\mathcal{O}) < \frac{2}{3}))$ if $n_{ele} < 1$ and $x_{ele} < 1$



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The SUSY breaking vacuum

Rank condition

$$F_{\mathcal{M}} = q_i^{\alpha} \tilde{q}_{\alpha}^j - \mu^2 \delta_i^j \qquad \alpha = 1, \dots, \widetilde{\mathcal{N}}, \quad i, j = 1, \dots, \mathcal{N}_1$$

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Since $N_1 > \widetilde{N} \to N_1 - \widetilde{N}$ SUSY breaking equations • $q = \widetilde{q}^T = \begin{pmatrix} \mu \mathbf{1}_{\widetilde{N}} \\ 0 \end{pmatrix} \quad p = \widetilde{p} = 0 \quad M = \begin{pmatrix} 0 & 0 \\ 0 & X \mathbf{1}_{N_1 - \widetilde{N}} \end{pmatrix}$ $L = \begin{pmatrix} 0 \\ Y \mathbf{1}_{N_1 - \widetilde{N}} \end{pmatrix} \quad K^T = \begin{pmatrix} 0 \\ \widetilde{Y} \mathbf{1}_{N_1 - \widetilde{N}} \end{pmatrix}$

• One loop CW for the pseudomoduli from W_{IR} :

 $W_{IR} = X\phi_1\phi_2 - \mu^2 X + \mu(\phi_1\phi_4 + \phi_2\phi_3) + Y\phi_1\phi_5 + Y\phi_2\phi_6 + \rho\phi_5\phi_6$

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- Assignment of R charges 0 or 2 for all the fields
- X, Y and \widetilde{Y} stable at the origin
- Possible sources of destabilizations
 - Corrections from the kahler potential
 - Non perturbative dynamics

Non perturbative dynamics and SUSY vacua

- At large M the fields q and \tilde{q} get a dynamical mass
- The fields p and \tilde{p} are massive at tree level
- Integrating out the massive fields there is a scale matching $\Lambda_L^{3\tilde{N}} = \tilde{\Lambda}_h^{3\tilde{N}-N_1-N_2} \det \rho \det M$

Non perturbative superpotential

$$W_{np} = ilde{N} \Lambda_L^3 = N \left(ilde{\Lambda}_h^{3 ilde{N} - N_1 - N_2} \det
ho \det M
ight)^{rac{1}{N_c}}$$

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SUSY vacua: all the fields at the origin except M

$$\langle M \rangle_{SUSY} = \mu \left(\frac{\mu}{\tilde{\Lambda}_h}\right)^{\frac{3\tilde{N}-N_1}{\tilde{N}}} \left(\frac{\tilde{\Lambda}_h}{\rho}\right)^{\frac{N_2}{\tilde{N}}}$$

The bounce action

•
$$S_B \sim \frac{M^4}{\mu^4} = \left(\frac{\mu}{\tilde{\Lambda}_h}\right)^4 \frac{3\tilde{N}-N_1}{\tilde{N}} \left(\frac{\tilde{\Lambda}_h}{\rho}\right)^{\frac{4N_2}{\tilde{N}}} \begin{vmatrix} S_B \text{ renormalized action} \\ \text{In the canonical basis} \\ \text{renormalization effects} \\ \text{on the couplings} \end{vmatrix}$$

if

Large Lifetime OR Small *W_{np}* at the origin

$$rac{
ho}{ ilde{\Lambda}_h} \ll \left(rac{\mu}{\Lambda}
ight)^{rac{3 ilde{N}-N_1}{N_2}}$$

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• Relations at the CFT exit scale Λ_c

$$\Lambda_{c} = \langle M \rangle_{SUSY} \gg \mu_{IR}, \rho_{IR}$$

$$Indeed$$

$$\epsilon_{IR} \equiv \frac{\rho_{IR}}{\mu_{IR}} \ll 1$$

$$\mu_{IR} = \Lambda_{c}^{-\frac{8\pi^{2}}{g_{*}^{2}(2\tilde{N}-N_{2})}} \epsilon_{IR}^{\frac{N_{2}}{(2\tilde{N}-N_{2})}} \ll \Lambda_{c}$$

$$\rho_{IR} = \Lambda_{c}^{-\frac{8\pi^{2}}{g_{*}^{2}(2\tilde{N}-N_{2})}} \epsilon_{IR}^{\frac{2\tilde{N}}{(2\tilde{N}-N_{2})}} \ll \Lambda_{c}$$

• The renormalized bounce action at
$$\Lambda_C$$
 is

$$S_B = \frac{e^{\frac{32\pi^2}{g_*^2(2\bar{N}-N_1))}}}{e^{\frac{4N_2}{\bar{N}-N_2}}} \qquad (As \text{ ISS if } N_2 \to 0)$$

Parametric dependence on the small relevant deformations

- What constraints have to be imposed in the UV?
 - ρ_{UV} and μ_{UV} small s.t. the model is approximately a SCFT
 - Bound on ϵ_{UV} s.t. it runs to an allowed ϵ_{IR}
- In the canonical basis analysis based on the running of the couplings in W

$$\mu_{IR} = \mu_{UV} Z_M^{-1} \text{ where } Z_{\rho IR} = \rho_{UV} Z_p Z_{\tilde{p}}^{-\gamma_i}$$

$$Z_{\phi_i} = \left(\frac{\Lambda_{IR}}{\Lambda_{UV}}\right)^{-\gamma_i} \text{ where here } \Lambda_{IR} = \Lambda_c$$

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Anomalous dimensions

- Δ_i = Scaling dimension of $\phi_i \rightarrow \Delta_i = 1 + \gamma_i/2$
- ► In a SCFT $\Delta_i = \frac{3}{2}R_i$, whith R_i the exact *R*-charge
- R-exact at the fixed point is computed by a-maximization

$$a_{trial}(R) = rac{3}{32rac{1}{1}} \left(3 Tr R^3 - Tr R \right)$$

 TrR^3 Coeff.of the gauge anomaly TrR Coeff. of the grav. anomaly

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 $R=R_0+\sum F_i$ non anomalous combinations of the U(1)s

Constraints from anomalies and from W impose

$$\blacktriangleright \quad \tilde{N} + N_1(R[q] - 1) + N_2(R[p] - 1) = 0, \quad 2R[q] + R[N] = 2, \quad R[K] + R[p] + R[q] = 2$$

 $R[q] = R[\tilde{q}], R[p] = R[\tilde{p}]$ and R[K] = R[L] from the symmetries

• Define
$$n = \frac{N_1}{N_2}$$
 and $x = \frac{\tilde{N}}{N_2}$

$$R[p] = rac{n-x+y_{max}}{n}$$

 $R[M] = 2y_{max}$

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• By using
$$\gamma_i = 3R_i - 2$$

$$\epsilon_{IR} = \epsilon_{UV} \left(\frac{\Lambda_c}{E_{UV}}\right)^{\frac{3}{2n}(n-2x+y_{max}(2-n))} \ll 1$$

Is this inequality satisfied in the whole CFT region ?
 Does E_{UV}/Λ_c have to be large?

Allowed region for the existence of metastable vacua in the weakly coupled region of the conformal window



$$\epsilon_{UV} = 10^{-2}$$

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Allowed region for the existence of metastable vacua in the weakly coupled region of the conformal window



$$\epsilon_{UV} = 10^{-4}$$

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UV hyerarchies

- Metastable vacua in SSQCD exist only if $\rho_{UV} \ll \mu_{UV}$
- This constraint can be imposed by hand
- Alternatively a dynamical mechanism is necessary
- If $N_2 = 1$ mass term from an SP(0) stringy instanton



•
$$S_{inst} = \alpha p \tilde{p} \beta \rightarrow W_{inst} = \int d\alpha d\beta e^{S_{inst}} = \Lambda e^{-A} p \tilde{p}$$

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Extension to other $\mathcal{N} = 1$ SCFTs

- Other SCFTs can possess metastable vacua in the IR
- Along the lines of the SSQCD case it is possible to give a general recipe for building these models
- ► Take a *N* = 1 gauge theory with metastable vacua in the IR free magnetic window
- Consider the dual theory and add some massive quarks, s.t. the theory is in a conformal window
- In most of the cases other fields and interactions must be added, s.t. no mesons hit the unitary bound
- S_B has a parametrical dependence on the relevant couplings
- ▶ With the RG running look for the UV regimes of scales and ranks consistent with the requirement of large S_B

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- Genericity in $\mathcal{N} = 1$ SCFTs
- R-symmetry breaking by relevant, marginal or irrelevant deformations
- Gauge mediation and SU(5) GUT
- Suppressing flavour anarchy in single sector of SUSY breaking
 - Small anomalous dimensions (near the IR free regime)
 - Long running in the CFT \rightarrow Landau poles tipically below the unification scale

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a-maximization with more groups not straightforward