# How does the string in flat space-time know about curved space-time?

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Curved space in string theory

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#### Foreword

- This talk is based on the work done together with
  - Giuseppe D'Appollonio, Rodolfo Russo and Gabriele Veneziano, to appear.

## Plan of the talk

#### 1 Introduction

- 2 The large distance behaviour of the classical solution
- 3 The approach of Amati, Ciafaloni and Veneziano(ACV)
- 4 The classical deflection angle in brane background
- 5 Scattering of a closed string on a Dp brane: disk
- 6 Deflection angle from string theory
- 7 The annulus diagram
- 8 Conclusions and outlook
- 9 The absorption cross section from a D3 brane

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## Introduction

- String theory, as originally formulated, is a theory in flat Minkowski space-time.
- It contains in its spectrum a massless spin 2 particle that has all the properties of a graviton.
- It turns out that the low-energy string effective action that one derives from scattering of strings is a (super)gravity theory with string corrections.
- Curved space-time is not put by hand, as in GR, but emerges from string scattering amplitudes.
- ► At the perturbative level (g<sub>s</sub> ~ 0) string theory is only a theory of strings.
- If we take into account non-perturbative effects string theory contains additional *p*-dimensional states called D(irichlet)p branes.

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On the one hand, they are classical solutions of the low-energy 10-dim string effective action:

$$S = \frac{1}{2\kappa_{10}^2} \int d^{10}x \sqrt{-g} \left[ R - \frac{1}{2} \left( \nabla \phi \right)^2 - \frac{1}{2(p+2)!} e^{-a\phi} \left( F_{p+2} \right)^2 \right]$$

coupled to graviton, dilaton and RR (p+1)-form potential given by:

$$ds^{2} = [H(r)]^{2A} \left( \eta_{\alpha\beta} dx^{\alpha} dx^{\beta} \right) + [H(r)]^{2B} \left( \delta_{ij} dx^{i} dx^{j} \right)$$

with 
$$r^2 \equiv \delta_{ij} x^i x^j$$
 and  
 $e^{-\phi(x)} = [H(r)]^{\frac{p-3}{8}}$ ,  $C_{01...p}(x) = ([H(r)]^{-1} - 1)$ 

A, B and are equal to

$$A = -\frac{7-p}{16}$$
,  $B = \frac{p+1}{16}$ 

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• H(r) is an harmonic function given by

$$H(r) = 1 + \left(\frac{R}{r}\right)^{7-p}; R_p^{7-p} = \frac{2\kappa_{10}T_pN}{(7-p)\Omega_{8-p}} = \frac{g_sN(2\pi\sqrt{\alpha'})^{7-p}}{(7-p)\Omega_{8-p}}$$
  
$$\kappa_{10} = \frac{(2\pi)^{7/2}}{\sqrt{2}}g_s(\alpha')^2; T_p = \frac{\sqrt{\pi}}{(2\pi\sqrt{\alpha'})^{p-3}}; \Omega_q = \frac{2\pi^{\frac{q+1}{2}}}{\Gamma(\frac{q+1}{2})}$$

Mass per unit volume and RR charge

$$M_{p} = rac{T_{p}}{\kappa_{10}} \; N = rac{(2\pi\sqrt{lpha'})^{1-p}}{2\pilpha' g_{s}} \; N \;\;\;;\;\;\; \mu_{p} = \sqrt{2} T_{p} \; N$$

Non-perturbative in  $g_s$ .

The low-energy string effective action is an action in curved space-time and therefore the classical solution will inherit all the properties of a theory of (super)gravity.

- On the other hand, in string theory the Dp branes are characterized by having open strings attached to their (p + 1)-dim world-volume.
- Therefore, the open strings satisfy Dirichlet boundary conditions along the directions transverse to the world-volume of the Dp brane.
- In string theory a Dp brane is described by a closed string state, called the boundary state.
- Dp branes interact by exchanging open strings and therefore the lowest order interaction is given by the annulus diagram.
   [Polchinski, 1995]
- By open/closed string duality the annulus diagram satisfies the following identity:

$$-\mathrm{Tr}\log(L_0-a) = \int_0^\infty \frac{d\tau}{\tau} \mathrm{Tr}\left(e^{-2\pi\tau(L_0-a)}\right) = \langle B|D|B\rangle$$

that determines the boundary state

$$|B
angle \equiv rac{T_{
ho}}{2} N|B_X
angle|B_\psi
angle$$
;  $D = rac{lpha'}{4\pi} \int_{|z|<1} rac{d^2z}{|z|^2} z^{L_0-a} ar{z} \widetilde{L}_0-a$ 

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The bosonic part of the boundary state is equal to

$$|B_X
angle = \delta^{d-p-1}(\hat{q}^i - y^i) \left(\prod_{n=1}^{\infty} e^{-rac{1}{n}lpha_{-n}S\cdot\widetilde{lpha}_{-n}}
ight)|0
angle_{lpha}|0
angle_{lpha}|p=0
angle$$

$$S \equiv (\eta_{lphaeta}; -\delta_{ij})$$

- Using the boundary state and the vertex operators for open and closed strings one can compute any amplitude involving scattering of strings on the D branes.
- In particular, these amplitudes determine the structure of the Born-Infeld action.
- All calculations are done in flat space where we are able to quantize the string.
- On the other hand, the D branes are classical solutions of the low-energy string effective action in curved space.
- How are we going to recover the curved space properties of the Dp branes computing scattering of strings in flat space?
- This is what I am going to show in some example in this seminar.

The large distance behaviour of the classical solution [M. Frau, A. Lerda, I. Pesando, R. Russo, S. Sciuto and PDV, 1997]

Given the boundary state one can compute the large distance behaviour of the various fields of the classical solution as follows:

$$\langle \psi | D | B 
angle$$
 ;  $D = rac{lpha'}{4\pi} \int rac{d^2 z}{|z|^2} z^{L_0 - a} ar{z} ilde{L}_{0 - a}$ 

 $\langle\psi|$  is the string state corresponding to the field of the classical solution.

 Let us start by computing the expression for the generic NS-NS massless field which is given by

$$J^{\mu\nu} \equiv {}_{-1} \langle \widetilde{0}, \frac{k_{\perp}}{2} | {}_{-1} \langle 0, \frac{k_{\perp}}{2} | \psi^{\nu}_{1/2} \ \widetilde{\psi}^{\mu}_{1/2} | D | B \rangle_{NS} = - \frac{T_{p}}{2k_{\perp}^{2}} V_{p+1} S^{\nu\mu}$$

 Specifying the different polarizations corresponding to the various fields we get for the dilaton

$$\delta\phi = \frac{1}{\sqrt{8}} \left( \eta^{\mu\nu} - k^{\mu}\ell^{\nu} - k^{\nu}\ell^{\mu} \right) J_{\mu\nu} = \frac{3-p}{4\sqrt{2}} T_{\rho} \frac{V_{\rho+1}}{k_{\perp}^2}$$

For the graviton we get:

$$\delta h_{\mu\nu}(k) = \frac{1}{2} \left( J_{\mu\nu} + J_{\nu\mu} \right) - \frac{\delta \phi}{\sqrt{8}} \eta_{\mu\nu}$$
$$= T_{\rho} \frac{V_{\rho+1}}{k_{\perp}^2} \operatorname{diag} \left( -A, A \dots A, B \dots B \right) ,$$

where  $A = -\frac{7-p}{16}$  and  $B = \frac{p+1}{16}$ For the Kalb-Ramond field we get:

$$\delta B_{\mu\nu}(k) = \frac{1}{\sqrt{2}} \Big( J_{\mu\nu} - J_{\nu\mu} \Big) = 0$$

In the R-R sector we get instead

$$\delta C_{01\dots p}(k) \equiv \langle P_{01\dots p}^{(C)} | D | B \rangle_{\mathrm{R}} = \mp \mu_p \frac{V_{p+1}}{k_{\perp}^2} ; \quad T_p = \sqrt{2}\mu_p$$

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We can express the previous fields in configuration space using the following Fourier transform valid for p < 7</p>

$$\int d^{(p+1)} x \, d^{(9-p)} x \frac{\mathrm{e}^{i k_{\perp} \cdot x_{\perp}}}{(7-p) \, r^{7-p} \, \Omega_{8-p}} = \frac{V_{p+1}}{k_{\perp}^2} \; ,$$

We must rescale the various fields according to

$$\varphi = \sqrt{2}\kappa_{10}\phi$$
 ,  $\tilde{h}_{\mu\nu} = 2\kappa_{10}h_{\mu\nu}$  ,  $C_{01...p} = \sqrt{2}\kappa_{10}C_{01...p}$ 

We get the following large distance behaviour for the dilaton

$$\delta\varphi(r) = \frac{3-\rho}{4} \left(\frac{R_{\rho}}{r}\right)^{7-\rho}$$

For the graviton we get

$$\delta \tilde{h}_{\mu\nu}(r) = 2 \left(\frac{R_{\rho}}{r}\right)^{7-\rho} \operatorname{diag}\left(-A, \ldots A, B \ldots B\right)$$

For the RR field we get

$$\delta \mathcal{C}_{01\dots p} = \left(\frac{R_p}{r}\right)^{7-p}$$

- ▶ The previous equations reproduce exactly the behavior for  $r \to \infty$  of the metric, the dilaton and the R-R potential of the solution.
- The next to the leading behaviour is expected to come from the one-point function with two boundary states:

$$\mathcal{N}\sum_{lpha,eta}\langle B|\int d^2z_1 \ W(z_1,ar{z}_1) \ D \ |B
angle_{lpha,eta}$$

W is the vertex operator corresponding to the massless closed string.

The explicit calculation gave zero after the sum over the spin structures [R. Marotta, I. Pesando, PDV (1998), unpublished].

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- Two alternative ways of getting the classical solution.
- One is by solving the classical supergravity equations of motion with the Dp brane as a source.
- The other is by computing the one-point function for the closed string fields in an action that contains their interaction in the bulk and their interaction with the D brane:

$$\langle \Phi(x) \; \mathrm{e}^{i S_{bulk} + i S_{boundary}} \rangle$$

#### where

$$S_{boundary} = \int d^{p+1}x \ T_p \left[ -\eta_{lphaeta} h^{lphaeta} + rac{\mathbf{3} - p}{2\sqrt{2}} \phi + \sqrt{2}C_{01...p} 
ight]$$

- By explicit calculation one can reproduce the leading and the next to the leading behaviour of the classical solution.
- It is still not clear why the stringy calculation does not work as in field theory. One may need an off-shell extrapolation.

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#### The approach of Amati, Ciafaloni and Veneziano(ACV) [ACV, 1987; Sundborg, 1987]

The starting point of the ACV approach is the four graviton scattering amplitude on the sphere in the Regge limit (s = 4E<sup>2</sup> → ∞ and small t):

$$\frac{a_{\text{Tree}}}{s} = \frac{32\pi G_{\text{N}}}{\alpha' s} \frac{\Gamma(-\frac{\alpha'}{4}t)}{\Gamma(1+\frac{\alpha'}{4}t)} \left(\frac{\alpha'}{4}s\right)^{2+\frac{\alpha'}{2}t} e^{-i\pi\frac{\alpha'}{4}t} \Longrightarrow 8\pi G_{\text{N}}\frac{s}{(-t)}$$

At high energy its Fourier transform in the space of impact parameter exponentiates and one can compute the classical deflection angle for large impact parameter:

$$\Theta = \frac{\sqrt{\pi} \Gamma(\frac{D}{2})}{\Gamma(\frac{D-1}{2})} \left(\frac{r_s}{b}\right)^{D-3} \quad ; \quad r_s^{D-3} = \frac{16\pi G_N \sqrt{s}}{(D-2)\Omega_{D-2}}$$

*D* is the number of non-compact directions and  $r_s$  is the Schwarzschild radius.

Then there are classical next to the leading corrections in the large impact parameter and string corrections.

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- It corresponds to the deflection angle of a particle moving in the Aichelburg-SexI metric that can be seen as the boosted Schwarzschild metric.
- It is the metric created by a fast moving particle as seen from the other particle.
- In the case of a scattering on a Dp brane, there is a background metric, namely the metric created by the Dp brane.
- How can we see the effect of this metric in the scattering of a massless closed string on a Dp brane?

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## The classical deflection angle in brane background

- At the classical level we can compute the deflection angle of a massless probe moving in the metric created by a Dp brane.
- Consider a general metric of the kind:

$$ds^{2} \equiv g_{\mu\nu}(x)dx^{\mu}dx^{\nu} = -\alpha(r)dt^{2} + \beta(r)(dr^{2} + r^{2}d\theta^{2})$$

where we have neglected coordinates that are not involved for a geodesic in which only t, r and  $\theta$  vary.

The geodesic equations can be best derived from the action of a massless point-particle in this metric:

$$S = \frac{1}{2} \int \frac{d\tau}{e} \dot{x}^{\mu} \dot{x}^{\nu} g_{\mu\nu}(x) = \frac{1}{2} \int \frac{d\tau}{e} \left( -\dot{t}^2 \alpha(r) + \beta(r) \left( \dot{r}^2 + r^2 \dot{\theta}^2 \right) \right)$$

where *e* is the einbein to take care of the reparametrization invariance of the world line coordinate  $\tau$ .

The conjugate momenta are given by:

$$p_{t} \equiv \frac{\partial L}{\partial \dot{t}} = -\frac{\dot{t}\alpha}{e} ; \quad p_{r} \equiv \frac{\partial L}{\partial \dot{r}} = \frac{\beta(\rho)\dot{r}}{e} ; \quad p_{\theta} = \frac{\partial L}{\partial \dot{\theta}} = \frac{\dot{\theta}r^{2}\beta(r)}{e}$$

The Eq. of motion for *e* gives:

$$\beta(\mathbf{r})\dot{\mathbf{r}}^{2} + \beta(\mathbf{r})\mathbf{r}^{2}\dot{\theta}^{2} = \alpha(\mathbf{r})\dot{t}^{2}$$

Since the Lagrangian does not depend explicitly on either t or θ there are two conserved quantities: the energy and the angular momentum

$$E = -lpha(r)\dot{t}$$
 ;  $J = eta(r)r^2\dot{ heta}$ 

where a dot denotes derivative with respect to  $\tau$  and we have taken e = 1.

Combining the three previous equations we get

$$\frac{\dot{\theta}}{\dot{r}} = \frac{J}{\beta r^2} \frac{1}{\sqrt{\frac{E^2}{\alpha\beta} - \frac{J^2}{\beta^2 r^2}}} = \frac{b}{r^2} \frac{1}{\sqrt{\frac{\beta}{\alpha} - \frac{b^2}{r^2}}}$$

where  $b \equiv J/E$  is the impact parameter.

The deflection angle is therefore given by:

$$\Theta_{p} = 2 \int_{r_{*}}^{\infty} \frac{dr}{r^{2}} \frac{b}{\sqrt{\frac{\beta}{\alpha} - \frac{b^{2}}{r^{2}}}} - \pi$$

r\* is the turning point i.e. the largest root of the equation  $\frac{\beta}{\alpha} - \frac{b^2}{r^2} = 0.$ 

- The result depends only on  $\alpha/\beta$ .
- It is therefore invariant under a *r*-dependent rescaling of the whole metric.
- Therefore, we can work alternatively in either the string or the Einstein frame.
- ▶ In our case we find, for a D*p* brane:

$$\frac{\beta}{\alpha} = 1 + \left(\frac{R_p}{r}\right)^{7-p}$$

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• Changing variable to  $u = \frac{b}{r}$  one gets:

$$\Theta_{p} = 2 \int_{0}^{u^{*}} \frac{du}{\sqrt{1 - u^{2} + \left(\frac{R_{p}}{b}\right)^{7-p} u^{7-p}}} - \pi$$

where  $u^*$  is the smallest root of the equation:

$$1-(u^*)^2+\left(rac{R_p}{b}
ight)^{7-p}(u^*)^{7-p}=0$$

• The integral can be done exactly for the cases p = 5, 6:

$$\tan \frac{\Theta_6}{2} = \frac{R_6}{2b} \quad ; \quad \Theta_5 = \frac{\pi}{\sqrt{1 - \left(\frac{R_5}{b}\right)^2}} - \pi$$

▶ For the case *p* = 3 we get instead:

$$\Theta_3 = 2\sqrt{1+k^2}K(k) - \pi$$
;  $K(k) = \int_0^1 \frac{dv}{\sqrt{(1-v^2)(1-k^2v^2)}}$ 

K is the complete elliptic integral of first kind.

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- For general p we have not yet been able to write the deflection angle in closed form.
- We have computed the leading and the next to the leading behaviour for large impact parameter:

$$\Theta_{\rho} = \sqrt{\pi} \left[ \frac{\Gamma(\frac{8-\rho}{2})}{\Gamma(\frac{7-\rho}{2})} \left( \frac{R_{\rho}}{b} \right)^{7-\rho} + \frac{1}{2} \frac{\Gamma(\frac{15-2\rho}{2})}{\Gamma(6-\rho)} \left( \frac{R_{\rho}}{b} \right)^{2(7-\rho)} + \dots \right]$$

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## Scattering of a closed string on a Dp brane: disk

- We consider the scattering of a massless closed string of the NS-NS sector on a Dp brane.
- ▶ The two closed strings have respectively momentum *p*<sub>1</sub> and *p*<sub>2</sub>.
- Along the directions of the world-volume of a Dp brane there is conservation of energy and momentum:

$$(p_1 + p_2)_{\parallel} = 0$$
 ;  $p_1^2 = p_2^2 = 0$ 

The scattering is described by two Mandelstam variables:

$$t = -(p_{1\perp} + p_{2\perp})^2 = -4E^2\cos^2rac{ heta}{2}$$
;  $s = E^2 = |p_{1\perp}|^2 = |p_{2\perp}|^2$ 

 $\theta$  = the angle between the d-dim ( $d \equiv 9 - p$ ) vectors  $p_{1\perp}$  and  $p_{2\perp}$ .

At high energy we consider the following kinematical configuration:

$$p_1 = (E, \underbrace{0 \dots 0}; E, \vec{p}_1)$$
;  $p_2 = (-E, \underbrace{0 \dots 0}; -E, \vec{p}_2)$ 

 $\vec{p}_1, \vec{p}_2$  are (d-1)-dim vectors orthogonal to the (p+1) direction.

At high energy  $(s \to \infty)$  only the coefficient of the term  $\text{Tr}(\epsilon_1 \epsilon_2^T)$  survives:

$$A_{1} = -\left(\frac{\kappa_{10}T_{\rho}N}{2}\right)(\alpha' E^{2})^{2}\frac{\Gamma(-\alpha' E^{2})\Gamma(-\frac{\alpha'}{4}t)}{\Gamma(1-\alpha' E^{2}-\frac{\alpha'}{4}t)}$$

where N is the number of D branes,

$$\kappa_{10} = \frac{(2\pi)^{7/2}}{\sqrt{2}} g_s(\alpha')^2 \; ; \; T_p = \frac{\sqrt{\pi}}{(2\pi\sqrt{\alpha'})^{p-3}} \; ; \; \frac{\kappa_{10}T_pN}{2} = \frac{R_p^{7-p}\pi^{\frac{9-p}{2}}}{\Gamma(\frac{7-p}{2})}$$

[ Ademollo et al, 1974, Klebanov and Thorlacius, 1995; Klebanov and Hashimoto, 1996, Garousi and Myers, 1996]

The poles in the t-channel correspond to exchanges of closed strings, while those in the s-channel correspond to exchanges of open strings:

$$2 + \frac{\alpha'}{2}t = 2m$$
;  $m = 2, 4, ...$ ;  $1 + \alpha' E^2 = n$ ;  $n = 1, 2...$ 

Regge behaviour at high energy:

$$A_{1} = \left(\frac{\kappa_{10}T_{\rho}N}{2}\right)e^{-i\frac{\alpha'}{4}t} \left(\sqrt{\alpha'}E\right)^{2+\frac{\alpha'}{2}t}\Gamma(-\frac{\alpha'}{4}t)$$

- A₁ diverges when E → ∞ and this creates problems with the unitarity of the S matrix.
- This problem is cured by higher orders in the perturbative expansion.
- The properly normalized S matrix is:

$$S = 1 + iT = 1 + irac{A}{\sqrt{2E_1}\sqrt{2E_2}} = 1 + irac{A}{2E}$$
;  $E_1 = E_2 \equiv E_2$ 

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The quantity that exponentiates at high energy and small t is:

$$iT_{1} \equiv i\frac{A_{1}}{2E} \Longrightarrow i\left(\frac{\kappa_{10}T_{p}N}{2}\right)e^{-j\frac{\alpha'}{4}t} \frac{(\sqrt{\alpha'}E)^{2+\frac{\alpha'}{2}t}}{2E}\Gamma(-\frac{\alpha'}{4}t)$$

Assume that the amplitude is dominated by the graviton massless pole at t ~ 0 (α' → 0):

$$iT_{1}(t, E) = i\left(\frac{\kappa_{10}T_{\rho}N}{2}\right)\left(\frac{2E}{(-t)} + i\pi\frac{\alpha' E}{2}\left(\sqrt{\alpha'}E\right)^{\frac{\alpha'}{2}t}\right)$$

- The real part describes the scattering of the closed string on the Dp brane, while the imaginary part describes the absorption of the closed string by the Dp brane.
- Go to impact parameter space:

$$T_1^R(b, E) + iT_1^I(b, E) = \int \frac{d^{d-1}q_t}{(2\pi)^{d-1}} e^{-ib \cdot q_t} T_1(t = -q_t^2, E)$$

## Deflection angle from string theory

For the real part one gets:

$$iT_1^R(b, E) \equiv 2i\delta(E, b) = i \frac{E \sqrt{\pi} R_p^{d-2}}{(d-3)b^{d-3}} \cdot \frac{\Gamma(\frac{d-1}{2})}{\Gamma(\frac{d-2}{2})}$$
;  $d \equiv 9 - p$ 

Assuming that it exponentiates, we get:

$$S(E,b) \equiv \mathrm{e}^{2i\delta(E,b)} = \mathrm{e}^{irac{E\sqrt{\pi}}{(d-3)b^{d-3}}\cdotrac{\Gamma(rac{d-1}{2})}{\Gamma(rac{d-2}{2})}}$$

Going back to momentum space, we get:

$$\int d^{d-1}b e^{i(b \cdot q_t + 2\delta(E,b))}$$

For large impact parameter we have the saddle point equation:

$$\vec{q_t} - \vec{b} \; \frac{E \; \sqrt{\pi} \; R_p^{d-2}}{b^{d-1}} \cdot \frac{\Gamma(\frac{d-1}{2})}{\Gamma(\frac{d-2}{2})} = 0$$

From which we compute the deflection angle:

$$\Theta_{p} = \frac{|q_{t}|}{E} = \sqrt{\pi} \left(\frac{R_{p}}{b}\right)^{7-p} \cdot \frac{\Gamma(\frac{8-p}{2})}{\Gamma(\frac{7-p}{2})}$$

It agrees with the classical calculation for large impact parameter!!

Assuming that also the imaginary part exponentiates, we get the absorption amplitude:

$$S^{abs}(E,b) = e^{-g_s \sqrt{\alpha'} E rac{(2\pi)^{rac{d-1}{2}}}{16} (\log(\sqrt{\alpha'}E))^{rac{1-d}{2}} e^{-rac{b^2}{2\alpha' \log(\sqrt{\alpha'}E)}}$$

that is a purely stringy effect.

To check the exponentiation and to compute the next to the leading behaviour in the expansion for large impact parameter we need to compute the annulus diagram.

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#### The annulus diagram

The annulus diagram is given by:

$$A_2 = \mathcal{N} \int d^2 z_a d^2 z_b \sum_{lpha,eta} \langle B | W_a(z_a, ar{z}_a) W_b(z_b, ar{z}_b) D | B 
angle_{lpha,eta}$$

 $W_{a,b}(z_{a,b}, \bar{z}_{a,b})$  are the closed string vertices,  $\mathcal{N}$  is a normalization factor and  $\sum_{\alpha,\beta}$  is the sum over the spin structures.

- The sum over the spin structures can be explicitly performed obtaining in practice only the contribution of the bosonic degrees of freedom without the bosonic partition function.
- The final result is rather explicit. [Pasquinucci, 1997 and Lee and Rey, 1997]

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► In the closed string channel the coefficient of the term with  $Tr(\epsilon_1 \epsilon_2^T)$  (relevant at high energy) of annulus diagram is equal to:

$$A_{2} = \left(\frac{\kappa_{10}}{\pi}\right)^{2} (8\pi^{2}\alpha')^{-\frac{p+1}{2}} \frac{(2\pi)^{4}}{2} N^{2} (\alpha's)^{2}$$

$$\times \int_{0}^{\infty} \frac{d\lambda}{\lambda} \lambda^{\frac{p+1}{2}} \int_{0}^{\frac{1}{2}} d\rho_{a} \int_{0}^{\frac{1}{2}} d\rho_{b} \int_{0}^{\frac{1}{\lambda}} d\omega_{a} \int_{0}^{\frac{1}{\lambda}} d\omega_{b}$$

$$\times e^{-\alpha'sV_{s}-\frac{\alpha'}{4}tV_{t}} ; \quad z_{a,b} \equiv e^{2\pi i(\rho_{a,b}+i\omega_{a,b})}$$

#### where

$$V_{s} = -2\pi\lambda\rho_{ab}^{2} + \log\frac{\Theta_{1}(i\lambda(\zeta + \rho_{ab})|i\lambda)\Theta_{1}(i\lambda(\zeta - \rho_{ab})|i\lambda)}{\Theta_{1}(i\lambda(\zeta + i\omega_{ab})|i\lambda)\Theta_{1}(i\lambda(\zeta - i\omega_{ab})|i\lambda)}$$

#### and

$$V_{t} = 8\pi\lambda\rho_{a}\rho_{b} + \log\frac{\Theta_{1}(i\lambda(\rho_{ab} + i\omega_{ab})|i\lambda)\Theta_{1}(i\lambda(\rho_{ab} - i\omega_{ab})|i\lambda)}{\Theta_{1}(i\lambda(\zeta + i\omega_{ab})|i\lambda)\Theta_{1}(i\lambda(\zeta - i\omega_{ab})|i\lambda)}$$

 $\rho_{ab} \equiv \rho_a - \rho_b \ ; \ \zeta = \rho_a + \rho_b \ ; \ \omega_{ab} \equiv \omega_a - \omega_b$ 

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- The high energy behaviour  $(E \rightarrow \infty)$  of the annulus diagram can be studied, by the saddle point technique, looking for points where  $V_s$  vanishes.
- This happens for  $\lambda \to \infty$  and  $\rho_{ab} \to 0$ .
- Performing the calculation one gets the leading term for  $E \to \infty$ :

$$A_2(E,t) \rightarrow \frac{i}{4E} \int \frac{d^{d-1}k}{(2\pi)^{d-1}} A_1(E,t_1) A_1(E,t_2) V(t_1,t_2,t)$$

where

$$t_1 \equiv -(rac{q}{2}+k)^2$$
 ;  $t_2 \equiv -(rac{q}{2}-k)^2$  ;  $t=-q^2$ 

and

$$V(t_1, t_2, t) = \frac{\Gamma(1 + \frac{\alpha'}{2}(t_1 + t_2 - t))}{\Gamma^2(1 + \frac{\alpha'}{4}(t_1 + t_2 - t))} \Longrightarrow 1$$

in the field theory limit ( $\alpha' \rightarrow 0$ ).

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Going to impact parameter space:

$$A_2(E,b) = \int rac{d^{d-1}q}{(2\pi)^{d-1}} \mathrm{e}^{-ib\cdot q} A_2(E,-q^2) = rac{i}{4E} (A_1(E,b))^2$$

• In terms of the matrix  $T \equiv \frac{A}{2E}$ :

$$T_2(E,b) = rac{i}{2} (T_1(E,b))^2$$

This implies that:

$$S(E,b) = 1 + iT_1 + iT_2 + \dots = 1 + iT_1 - \frac{1}{2}(T_1)^2 + \dots = e^{iT_1}$$

- At high energy the amplitude exponentiates: no problems with unitarity.
- ► We have extracted the leading behaviour at high energy:

$$T_1 \sim E$$
 ;  $T_2 \sim E^2$ 

But there is also a next to the leading contribution to  $\frac{T_2^{nl}}{2} \sim E$ .

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- It can be extracted from the annulus amplitude.
- It must also exponentiate.
- It gives the next to the leading correction to the deflection angle that agrees with the one obtained from the classical calculation.
- We have seen that both the leading and the next to the leading terms in the expansion for large impact parameter are reproduced by the string scattering amplitude.
- Being brave, one can assume that all the terms are reproduced.
- This implies an equation for the phase shift:

$$-\frac{2}{E}\frac{\partial\delta(E,b)}{\partial b}=\Theta+\pi$$

Integrating the previous equation one gets, for the simple case p = 5:

$$\delta_5(E,b) = \frac{\pi E b}{2} \left[ 1 - \sqrt{1 - \left(\frac{R_5}{b}\right)^2} \right]$$

- The phase shift has an imaginary part for  $b \le R_5$ .
- For b ≤ R<sub>5</sub> the probe particle is absorbed by the brane with an absorption cross section that is proportional to:

$$|\mathcal{S}| \sim e^{-2\pi \textit{ER}_5}$$

and is independent from b.

Similar results for any *p*.

## Conclusion and outlook

- We have seen how from string scattering in flat space-time we can recover properties of curved space-time.
- In particular, from the scattering of a massless closed string on a Dp brane at high energy and low transfer momentum we have computed the deflection angle of a probe particle moving in the metric of the Dp brane.
- The result reproduces the leading and the next to the leading contributions for large impact parameter computed from classical gravity in the metric of a Dp brane.
- String corrections to the field theory results can also be computed.
- We have not seen any effect from the dilaton and the RR field on the deflection angle.
- ► Work directly with the angular momentum *J* instead of the impact parameter  $b (b \sim \frac{J}{E})$ .

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## The absorption cross section from a D3 brane [Klebanov, 1997]

The low energy absorption cross section of a dilaton by a D3 brane can be calculated and one gets:

$$\sigma_{abs.D3} = rac{\pi^4}{8} \omega^3 R^8$$
 ;  $H(r) = 1 + rac{R^4}{r^4}$ 

 $\omega$  is the dilaton energy.

- This calculation is done using the curved space formalism in the metric of a D3 brane.
- On the other hand, the same quantity can be computed from the the scattering of a closed string on a D3 brane that generates two open strings (gluons).
- At low energy the coupling of the DBI action that is relevant is the one involving a dilaton and two gauge fields that gives the following amplitude:

$$\mathcal{A} = -\frac{\kappa_{10}}{\sqrt{2}} \cdot 2 \cdot \frac{p_1 \cdot p_2}{\sqrt{2\omega} \omega} = -\frac{\kappa_{10}\sqrt{\omega}}{\sqrt{2}} \quad ; \quad E_1 = E_2 = \frac{\omega}{2}$$

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The absorption cross section is equal to:

$$\frac{1}{2}\int \frac{d^3p_1}{(2\pi)^3}\int \frac{d^3p_2}{(2\pi)^3} (2\pi)^4 \delta(E_1+E_2-\omega)\delta^3(\vec{p}_1+\vec{p}_2)\mathcal{A}^2$$

Factor <sup>1</sup>/<sub>2</sub> because the two particles in the final state are identical.
One gets:

$$\sigma_{abs.D3} = \frac{\kappa_{10}^2 N^2 \omega^3}{32\pi} = \frac{\pi^4}{8} \omega^3 R^8 \quad ; \quad R^4 = \frac{\kappa_{10} N}{2\pi^{\frac{5}{2}}}$$

- The same result from the coupling of a dilaton with two gauge fields that has apriori nothing to do with curved space-time.
- ► This calculation is at the origin of the Maldacena conjecture.

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