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D3-D7 Quark-Gluon Plasmas

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based on works with

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Plan

- Motivations: quark-gluon plasma @ RHIC
- D3-D7 plasmas
- Thermodynamics and hydrodynamics
- Novel universal results

Motivations: Heavy ion collisions at RHIC



[3000 particles created in a central Au+Au collisions at RHIC (STAR)]



 Δ t = 10 fm/c = 3.3.10⁻²³ s

[Westfall 2006]

- Au-Au collision at Relativistic Heavy Ion Collider (RHIC)
- At T>Tc a plasma of quarks and gluons.
- Elliptic flow, jet quenching... : QGP probably strongly coupled
- Liquid with very low viscosity. Model by relativistic hydrodynamics
- A challenge for theoretical physics: pQCD not very useful

• Thermodynamics (equilibrium) properties from Lattice QCD



- At T>Tc, small T dependence (nearly conformal)
- At T>Tc , 80% of free Stefan-Boltzmann gas of quarks and gluon
- What about hydrodynamics? On lattice problems with real-time

- QGP@RHIC nearly conformal, strongly coupled
- $\eta/s \le 4(1/4\pi)$ (hydro simulations, elliptic flow)
- $\epsilon/\epsilon_0 \sim 80\%$ (lattice QCD)
- Thermal CFTs (e.g. N=4 SYM) with AdS5-BH dual as benchmarks!
- $\eta/s = 1/4\pi$ [Policastro, Son, Starinets 2001]
- $\epsilon/\epsilon_0 = s/s_0 = 3/4$ [Witten 1998]
- Using gauge/gravity duality, we can study real-time processes
- Focus on universal properties (e.g. $\eta/s=(1/4\pi)$ for every plasma with gravity dual [Kovtun, Son, Starinets 2004])

Prototypes: planar strongly coupled thermal quivers on Nc D3branes at CY3 cones.

 \mathcal{N} =1 superconformal theories at T=0

- Dual description (T≠0) : IIB on AdS5 (black hole) x X5, constant dilaton and F5 RR flux. X5: Sasaki-Einstein base of the cone.
- λ= g² Nc >> 1, Nc>>1
- Matter fields in 2-index representation (adjoint, bifundamental...): no quarks (i.e fundamental matter fields)!
- Goal: add dynamical flavors and study their effects on the properties of the strongly coupled plasmas

- Flavors can be added by means of D7-branes, along 4d Minkowski and wrapped on non-compact 4-manifold.
- To account for vacuum polarization effects due to dynamical flavors, i.e. to go beyond the so-called quenched approximation, we need to account for the backreaction of the flavor branes on the background.
- Not an easy task if flavor branes are localized, on top of each other. Delta function sources and p.d.e.
- Most of the known results at finite T concern the quenched approximation where the flavor branes are treated as probes [Karch-Katz 02].

Technology

 Take Nf >>1 D7-branes homogeneously smeared over transverse space. [F.B., Casero, Cotrone, Kiritsis, Paredes 05] This simplifies sugra:

1) delta function sources $\rightarrow \Omega$ density distribution 2-form

2) partial diff. e.o.m \rightarrow ordinary diff. e.o.m in a radial variable

Preserve most of the isometries of unflavored background

- Consider a perturbative expansion in $\epsilon \sim \lambda Nf/Nc=g^2Nf$. Up to second order this allows to find analytic solutions.
- D7 induce dilaton blow up: UV Landau pole. Decouple IR: cutoff $rh \le r < r^* < < rLP$, i.e. $T < < \Lambda_{LP}$. Ensured taking $\epsilon < < 1$

Dual field theory. Example: D3-D7 in flat space. Zi, i=1,2,3 complex transverse coordinates Karch-Katz embedding for D7: Z1=0 (breaks SU(4)R) Superpotential $W = \Phi_1[\Phi_2, \Phi_3] + \tilde{q}\Phi_1 q$. N=4 \rightarrow N=2 susy

Conformality broken at quantum level (massless flavors) UV Landau Pole

 $\begin{array}{ll} \mbox{Generalized embedding} & \sum_1^3 a_i Z_i = 0, & \sum_1^3 |a_i|^2 = 1. \\ \mbox{Superpotential} & W = \Phi_1[\Phi_2, \Phi_3] + \tilde{q}(a_1 \Phi_1 + a_2 \Phi_2 + a_3 \Phi_3)q \end{array}$

Smearing: Nf>>1 such terms, integrate over ai, recovers SU(3)x U(1), preserves N=1 susy. Flavor group is U(1)^(Nf)

 $S = S_{IIB} + S_{fl}$

$$S_{IIB} = \frac{1}{2\kappa_{10}^2} \int d^{10}x \sqrt{-g_{10}} \left[R - \frac{1}{2} \partial_M \Phi \partial^M \Phi - \frac{1}{2} e^{2\Phi} F_{(1)}^2 - \frac{1}{2} \frac{1}{5!} F_{(5)}^2 \right]$$

$$S_{fl} = -T_7 \sum_{N_f} \int_{\mathcal{M}_8} d^8 \xi \, e^{\Phi} \sqrt{-\det \hat{g}_8} \, + \, T_7 \sum_{N_f} \int_{\mathcal{M}_8} \hat{C}_8 \tag{PDEs}$$

$$S_{fl}^{smear} = -T_7 \int_{\mathcal{M}_{10}} d^{10}x \sqrt{-\det g_{10}} e^{\Phi} |\Omega| + T_7 \int_{\mathcal{M}_{10}} \Omega \wedge C_8 \quad \text{(ODEs)}$$

$$dF_{(1)} = -g_s \Omega \qquad \qquad \Omega \equiv -g_s \sum_i \Omega^{(i)}, \quad |\Omega| \equiv \sum_i \sqrt{\frac{1}{2} \Omega_{MN}^{(i)} \Omega_{PQ}^{(i)} g^{MN} g^{PQ}}$$

[X5 Sasaki-Einstein: $ds_{X_5}^2 = ds_{KE}^2 + (d\tau + A_{KE})^2$, $dA_{KE} = 2J_{KE}$]

$$ds_{10}^2 = h^{-\frac{1}{2}} \left[-b \, dt^2 + d\vec{x}_3^2 \right] + h^{\frac{1}{2}} \left[S^8 F^2 b^{-1} \, dr^2 + r^2 ds_5^2 \right]$$
$$h = \frac{R^4}{r^4}, \quad b = 1 - \frac{r_h^4}{r^4}$$

$$ds_5^2 = S^2 ds_{KE}^2 + F^2 (d\tau + A_{KE})^2, \quad dA_{KE} = 2J_{KE}$$

$$\Phi = \Phi(r), \quad F_{(5)} = Q_c (1 + *) \operatorname{vol}(X_5)$$

$$F_{(1)} = Q_f (d\tau + A_{KE}), \quad dF_{(1)} = 2Q_f J_{KE} \equiv -g_s \Omega$$

$$R^{4} = \frac{Q_{c}}{4}, \quad Q_{c} = \frac{(2\pi)^{4} g_{s} N_{c} \, \alpha'^{2}}{Vol(X_{5})}, \quad Q_{f} = \frac{g_{s} N_{f} Vol(X_{3})}{4 Vol(X_{5})}$$

$$F = 1 - \frac{\epsilon_*}{24} \left(1 + \frac{2r^4 - r_h^4}{6r_*^4 - 3r_h^4}\right)$$

$$S = 1 + \frac{\epsilon_*}{24} \left(1 - \frac{2r^4 - r_h^4}{6r_*^4 - 3r_h^4}\right)$$

$$\Phi = \Phi_* + \epsilon_* \log \frac{r}{r_*}$$

 $\epsilon \equiv \mathsf{Qf} \exp (\Phi)$

$$\begin{split} F &= 1 - \frac{\epsilon_*}{24} \left(1 + \frac{2r^4 - r_h^4}{6r_*^4 - 3r_h^4} \right) + \frac{\epsilon_*^2}{1152} \left(17 - \frac{94}{9} \frac{2r^4 - r_h^4}{2r_*^4 - r_h^4} + \frac{5}{9} \frac{(2r^4 - r_h^4)^2}{(2r_*^4 - r_h^4)^2} + \right. \\ &\left. - \frac{8}{9} \frac{r_h^8(r_*^4 - r^4)}{(2r_*^4 - r_h^4)^3} - 48 \log(\frac{r}{r_*}) \right) + O(\epsilon_*^3) \;, \\ S &= 1 + \frac{\epsilon_*}{24} \left(1 - \frac{2r^4 - r_h^4}{6r_*^4 - 3r_h^4} \right) + \frac{\epsilon_*^2}{1152} \left(9 - \frac{106}{9} \frac{2r^4 - r_h^4}{2r_*^4 - r_h^4} + \frac{5}{9} \frac{(2r^4 - r_h^4)^2}{(2r_*^4 - r_h^4)^2} + \right. \\ &\left. - \frac{8}{9} \frac{r_h^8(r_*^4 - r^4)}{(2r_*^4 - r_h^4)^3} + 48 \log(\frac{r}{r_*}) \right) + O(\epsilon_*^3) \;, \\ \Phi &= \Phi_* + \epsilon_* \log \frac{r}{r_*} + \frac{\epsilon_*^2}{72} \left(1 - \frac{2r^4 - r_h^4}{2r_*^4 - r_h^4} + 12 \log \frac{r}{r_*} + 36 \log^2 \frac{r}{r_*} + \right. \\ &\left. + \frac{9}{2} \left(Li_2(1 - \frac{r_h^4}{r^4}) - Li_2(1 - \frac{r_h^4}{r_*^4}) \right) \right) + O(\epsilon_*^3) \end{split}$$

Focus on deep IR physics T<< Λ *<< Λ LP

$$F = 1 - \frac{\epsilon_h}{24} + \frac{17}{1152}\epsilon_h^2 - \frac{\epsilon_h^2}{24}\log\frac{r}{r_h} \equiv F_h - \frac{\epsilon_h^2}{24}\log\frac{r}{r_h},$$

$$S = 1 + \frac{\epsilon_h}{24} + \frac{1}{128}\epsilon_h^2 + \frac{\epsilon_h^2}{24}\log\frac{r}{r_h} \equiv S_h + \frac{\epsilon_h^2}{24}\log\frac{r}{r_h},$$

$$\Phi = \Phi_h + \epsilon_h\log\frac{r}{r_h} + \frac{\epsilon_h^2}{6}\log\frac{r}{r_h} + \frac{\epsilon_h^2}{2}\log^2\frac{r}{r_h} + \frac{\epsilon_h^2}{16}Li_2\left(1 - \frac{r_h^4}{r^4}\right)$$

$$\epsilon_h = \epsilon_* + \epsilon_*^2 \log \frac{r_h}{r_*}, \qquad T \frac{d\epsilon_h}{dT} = \epsilon_h^2$$

- Massless-flavored susy solution (b=1) exactly known. It has a dilaton blowing up at rLP (UV Landau pole) and a (good) singularity in the IR [Benini, Canoura, Cremonesi, Nunez, Ramallo 06].
- Non extremal solution is regular at the horizon
- D7-brane distribution dense enough if Nf>>1
- Regime of validity

Nc>>1, λ >>1, Nf>>1, $\epsilon \sim \lambda Nf/Nc <<1$

Thermodynamics

 $\epsilon_h \equiv \frac{\lambda_h \operatorname{Vol}(X_3)}{16\pi \operatorname{Vol}(X_5)} \frac{N_f}{N_c} \qquad \qquad \frac{d\epsilon_h}{dT} = \frac{\epsilon_h^2}{T} + O(\epsilon_h^3)$

$$s = \frac{2\pi A_8}{\kappa_{(10)}^2 V_3} = \frac{\pi^5}{2Vol(X_5)} N_c^2 T^3 \left[1 + \frac{1}{2} \epsilon_h + \frac{7}{24} \epsilon_h^2 + O(\epsilon_h^3) \right]$$

$$\varepsilon = \frac{E_{ADM}}{V_3} = \frac{3}{8} \frac{\pi^5}{Vol(X_5)} N_c^2 T^4 \left[1 + \frac{1}{2} \epsilon_h(T) + \frac{1}{3} \epsilon_h(T)^2 + O(\epsilon_h(T)^3) \right]$$

$$\frac{F}{V_3} = -p = \varepsilon - Ts = -\frac{1}{8} \frac{\pi^5}{Vol(X_5)} N_c^2 T^4 \left[1 + \frac{1}{2} \epsilon_h(T) + \frac{1}{6} \epsilon_h(T)^2 + O(\epsilon_h(T)^3) \right]$$

Check: first order results match with those obtained in the literature in the probe approximation [Mateos, Myers, Thomson 07]

Thermodynamics

Breaking of conformal invariance: a second order effect

Interaction measure (trace of stress energy tensor)

$$(\varepsilon - 3p)/T^4 = [\pi^5 N_c^2 / 16 Vol(X_5)] \epsilon_h(T)^2$$

Speed of sound

$$v_s^2 = \frac{s}{c_V} = \frac{1}{3} \left[1 - \frac{1}{6} \epsilon_h(T)^2 + O(\epsilon_h(T)^3) \right]$$

Expect non zero bulk viscosity at second order

Hydrodynamics

- Effective theory of long wavelength fluctuations around local thermal equilibrium.
- Zero charge second order relativistic hydro

[Baier,Romatschke,Son,Starinets,Stephanov 2008; Romatschke 2009; Bhattacharyya, Hubeny, Minwalla, Rangamani 2008]

$$\begin{aligned} \nabla_{\mu}T^{\mu\nu} &= 0 & T^{\mu\nu} = \varepsilon u^{\mu}u^{\nu} + p\Delta^{\mu\nu} + \pi^{\mu\nu} + \Delta^{\mu\nu}\Pi \\ \varepsilon &= \text{energy density} \\ u^{\mu} &= \text{velocity field} \\ p(\varepsilon) &= \text{pressure} \\ \Delta^{\mu\nu} &= g^{\mu\nu} + u^{\mu}u^{\nu} \\ g^{\mu\nu} &= \text{metric} \end{aligned}$$
 Dissipative terms

Hydrodynamics

$$T^{\mu\nu} = \varepsilon u^{\mu} u^{\nu} + p \Delta^{\mu\nu} + \pi^{\mu\nu} + \Delta^{\mu\nu} \Pi$$

Expansion in gradients of u^{μ} and $\varepsilon \rightarrow s$ (via thermodynamics):

$$\begin{aligned} \pi^{\mu\nu} &= -\eta \sigma^{\mu\nu} + \eta \tau_{\pi} \Big[\langle D\sigma^{\mu\nu} \rangle + \frac{\nabla \cdot u}{3} \sigma^{\mu\nu} \Big] + \kappa \Big[R^{<\mu\nu>} - 2u_{\alpha}u_{\beta}R^{\alpha<\mu\nu>\beta} \Big] \\ &+ \lambda_{1}\sigma_{\lambda}^{<\mu}\sigma^{\nu>\lambda} + \lambda_{2}\sigma_{\lambda}^{<\mu}\Omega^{\nu>\lambda} + \lambda_{3}\Omega_{\lambda}^{<\mu}\Omega^{\nu>\lambda} + \kappa^{*}2u_{\alpha}u_{\beta}R^{\alpha<\mu\nu>\beta} \\ &+ \eta \tau_{\pi}^{*} \frac{\nabla \cdot u}{3} \sigma^{\mu\nu} + \lambda_{4}\nabla^{<\mu}\log s \nabla^{\nu>}\log s \end{aligned}$$

$$\Pi = -\zeta (\nabla \cdot u) + \zeta \tau_{\Pi} D (\nabla \cdot u) + \xi_1 \sigma^{\mu\nu} \sigma_{\mu\nu} + \xi_2 (\nabla \cdot u)^2 + \xi_3 \Omega^{\mu\nu} \Omega_{\mu\nu} + \xi_4 \nabla^{\perp}_{\mu} \log s \nabla^{\mu}_{\perp} \log s + \xi_5 R + \xi_6 u^{\alpha} u^{\beta} R_{\alpha\beta}$$

 $\eta =$ shear viscosity, $\zeta =$ bulk viscosity, $au_{\pi}, au_{\Pi} =$ relaxation times.

Holographic calculations

Holographically:

- Each fluid mode has dual gravity mode.
- Fluctuate corresponding mode (gravity equations).
- Extract dispersion relations (for hydrodynamic modes).
- Correlators from basic AdS/CFT prescription:

$$\langle e^{-\int \phi_0 \mathcal{O}} \rangle = e^{-S_{gravity}(\phi_0)}$$

(extended to real time case).

Holographic calculations

Standard calculation of (some) coefficients.

Fluctuations $\sim e^{-i(\omega t - qz)}h_{(\cdot,\cdot)} \Rightarrow$ modes classified by $SO(2)_{x,y}$.

From dispersion relation of scalar modes:

$$\omega = c_s q - i\Gamma q^2 + \frac{\Gamma}{c_s} \left(c_s^2 \tau^{eff} - \frac{\Gamma}{2} \right) q^3 + \mathcal{O}(q^4)$$
$$\Gamma = \frac{\eta}{sT} \left(\frac{2}{3} + \frac{\zeta}{2\eta} \right) \qquad \tau^{eff} = \frac{\tau_\pi + \frac{3\zeta}{4\eta} \tau_\Pi}{1 + \frac{3\zeta}{4\eta}}$$

From retarded correlator of tensorial mode:

$$G_R^{xy,xy} = p - i\eta\omega + \left(\frac{\eta\tau_\pi}{2} - \frac{\kappa}{2} + \kappa^*\right)\omega^2 - \frac{\kappa}{2}q^2 + \mathcal{O}(q^3,\omega^3)$$

Results [Bigazzi-Cotrone-Tarrio 2009]:

$$\begin{split} c_s &= \frac{1}{\sqrt{3}} \left(1 - \frac{1}{12} \epsilon_h^2 \right) \\ \frac{\eta}{s} &= \frac{1}{4\pi} \qquad [\text{Kovtun} - \text{Son} - \text{Starinets 2005}] \\ \frac{\zeta}{\eta} &= \frac{1}{9} \epsilon_h^2 \quad \Rightarrow \quad \text{Buchel's bound saturated at } \mathcal{O}(\epsilon_h^2). \text{ Small.} \\ \tau^{eff}T &= \tau_{\pi,0} T_0 + \frac{16 - \pi^2}{128\pi} \epsilon_h^2 \quad \Rightarrow \quad \text{Positive correction (bound?).} \\ \frac{T\kappa}{s} &= \frac{T_0 \kappa_0}{s_0} \left(1 - \frac{\epsilon_h^2}{8} \right) \quad \Rightarrow \quad \text{Negative correction (bound?).} \\ \frac{T^2}{p} (\kappa^* + \eta \tau_\pi) &= \frac{T_0^2}{p_0} \eta_0 \tau_{\pi,0} + \frac{T_0^2}{p_0} \eta_0 \left(\frac{\tau_{\pi,0}}{8} - \frac{16 + \pi^2}{128\pi T_0} \right) \epsilon_h^2 \end{split}$$

$$\tau_{\pi,0}T_0 = \frac{2 - \log 2}{2\pi} \qquad \frac{T_0\kappa_0}{s_0} = \frac{1}{4\pi^2} \qquad \frac{T_0\eta_0}{p_0} = \frac{1}{\pi}$$

Simpler derivation

5d reduction

$$S_5 = \frac{Vol(X_5)}{2\kappa_{10}^2} \int d^5x \sqrt{-\det g} \left[R[g] - \frac{40}{3} (\partial f)^2 - 20(\partial w)^2 - \frac{1}{2} (\partial \Phi)^2 - V(\Phi, f, w) \right]$$

$$V(\Phi, f, w) = 4e^{\frac{16}{3}f + 2w} \left(e^{10w} - 6 + Q_f e^{\Phi} \right) + \frac{1}{2}Q_f^2 e^{\frac{16}{3}f - 8w + 2\Phi} + \frac{Q_c^2}{2}e^{\frac{40}{3}f}$$

AdS vacuum at Qf=0 (i.e. ϵ h=0): f=w= Φ =0. Fluctuating around:

- $f \leftrightarrow \Delta = 8 \Rightarrow$ Irrelevant. $w \leftrightarrow \Delta = 6 \Rightarrow$ Irrelevant. $\Phi \leftrightarrow \Delta = 4 \Rightarrow$ Marginally Marginally irrelevant.

 Φ Breaks conformality at leading order (dual to flavor term in superpotential)

Simpler derivation

• At leading order only Φ contributes $V \sim -12 + 4 \text{ Qf } \exp(\Phi) + O(\epsilon h^2)$

$$\mathcal{V}_{\phi} \equiv \left(\frac{V'(\phi_h)}{V(\phi_h)}\right)^2 \qquad \qquad \mathcal{V}_{f,w} = 0, \quad \mathcal{V}_{\Phi} = \frac{\epsilon_h^2}{9}$$

- Note: $T d(\epsilon h^2)/dT = O(\epsilon h^3)$ i.e. it is a constant in our limits
- System effectively 5d gravity plus single scalar with exponential potential V=Vo e^(a Φ) [Chamblin-Reall model]
- For these models first order hydro coefficients simply given in terms of the potential [Gubser,Pufu,Rocha 2008]

$$\frac{\zeta}{\eta} = \frac{V'(\phi_h)^2}{V(\phi_h)^2} = \frac{\epsilon_h^2}{9} \,. \qquad c_s^2 - \frac{1}{3} = -\frac{1}{2} \frac{V'(\phi_h)^2}{V(\phi_h)^2} = -\epsilon_h^2/18$$

Simpler derivation

• In [Kanitscheider-Skenderis 2009]: Chamblin-Reall models

$$S = \frac{1}{2\kappa^2} \int d^5x \sqrt{-\det g} \left[R[g] - \frac{1}{2} (\partial\phi)^2 - V_0 e^{\gamma\phi} \right]$$

- Can be obtained from toroidal compactifications of pure gravity+cosmo constant in 2σ +1 dimensions and continuation in σ !
- $Cs^2 = 1/3 \gamma^2/2 = (higher dim CFT) = 1/(2\sigma-1)$
- For D3-D7, $\sigma = 2 + \epsilon h^2/4 + O(\epsilon h^3)$
- Crucial point: everything smooth in σ in 5d
- From known hydrodynamics on $AdS_{2\sigma+1}$ obtain hydro by dimensional reduction

[Romatschke 2009, Bigazzi, Cotrone, to appear]

All the transport coefficients :



[Saturation of Buchel's bound]

Summarizing

- Dual of a theory with conformality broken by marginally irrelevant operators (flavor terms) behaves effectively, at leading order, like a Chamblin-Reall model.
- Transport coefficients from [Kanitscheider-Skenderis 2009]
- Same holds for marginally relevant operators (e.g. cascading plasmas)

Generalizations

- QGP (RHIC, LHC)nearly conformal and strongly coupled for $1.5T_c \leq T \leq 4T_c$.
- Model it holographically with AdS5 (BH) deformed by a scalar dual to marginally relevant operator
- Use previous relations and estimate all the transport coefficients (to second order) at RHIC and LHC: just one parameter to fit, for example the speed of sound (for which we have quite solid lattice estimates)

Estimates for QCD

• From lattice [Katz 2005]: $c_s^2(T \sim 1.5T_c) \sim 0.283$ (RHIC)

$\frac{\eta}{s}$	$\frac{1}{4\pi}$	$T\tau_{\pi}$	0.222	$\frac{T\kappa}{s}$	0.022
$\frac{T\lambda_1}{s}$	0.014	$\frac{T\lambda_2}{s}$	-0.021	$\frac{T\lambda_3}{s}$	0
$\frac{T\kappa^*}{s}$	-0.006	$T au_{\pi}^{*}$	-0.031	$\frac{T\lambda_4}{s}$	0
$\frac{\zeta}{\eta}$	0.101	$T\tau_{\Pi}$	0.208	$\frac{T\xi_1}{s}$	0.001
$\frac{T\xi_2}{s}$	0.001	$\frac{T\xi_3}{s}$	0	$\frac{T\xi_4}{s}$	0
$\frac{T\xi_5}{s}$	0.001	$\frac{T\xi_6}{s}$	0.004		

Comments and future directions

- Studied jet quenching parameter (flavors enhance it)
- Equations extended to massive flavors
- Study of probes on these backgrounds and conductivity
- Finite baryon density (work in progress)
- Condensed matter applications
- Non-relativistic limits

Thank you

Simplest example: Nc D3 + Nf D7 in flat space (X5=S⁵) at T=0



•At Nf=0 a SCFT: N=4 SU(Nc) SYM (SU(4)R). In N=1 components:

 $W_0 = \Phi_1[\Phi_2, \Phi_3]$ $SU(3) \times U(1)_R$ symmetry.

Add Nf D7-branes wrapping non compact 4manif (ex. Z1 = µ1) D3-D7 strings↔ fundamental hypers. SU(Nf) flavor symmetry

 $W_2 = W_0 + \tilde{q}_i(\Phi_1 - m_1)q_i$

Break global symmetry, conformal invariance and susy $N=4 \rightarrow N=2$, b0 = (3Nc - 3Nc) - Nf \rightarrow UV Landau pole

We will consider N=1 setups. They will inherit this UV behavior. We will focus on IR physics well below the Landau Pole.

The smearing

Take Nf>>1 D7-branes homogeneously smeared over transverse space. Preserve original global symmetries and N=1 susy (at T=0).

Generalized embedding $\sum_{i=1}^{3} a_i Z_i = \mu$, $\sum_{i=1}^{3} |a_i|^2 = 1$

$$W = \Phi_1[\Phi_2, \Phi_3] + \tilde{q}(a_1\Phi_1 + a_2\Phi_2 + a_3\Phi_3 - m)q$$

Smearing= sum over flavors, integrate over ai in W. Recover SU(3). Flavor symmetry explicitly broken to U(1)^Nf.

Let's focus on massless case m=0 (i.e. μ =0): D7 reach the origin. Massless susy embeddings also solve D7 worldvolume equations in nonextremal case.

Transport coefficients of AdS_6 :

[Bhattacharyya-Loganayagam-Mandal-Minwalla-Sharma 2008]

$$\eta^{(2\tilde{\sigma})} = \frac{s^{(2\tilde{\sigma})}}{4\pi}$$

$$\kappa^{(2\tilde{\sigma})} = \frac{2\eta^{(2\tilde{\sigma})}2\tilde{\sigma}}{4\pi T(2\tilde{\sigma}-2)}$$

$$\tau^{(2\tilde{\sigma})}_{\pi} = \frac{2\tilde{\sigma}}{4\pi T} \left(1 - \int_{1}^{\infty} \frac{y^{2\tilde{\sigma}-2} - 1}{y(y^{2\tilde{\sigma}}-1)} dy\right)$$

$$\lambda_{1}^{(2\tilde{\sigma})} = \frac{\eta^{(2\tilde{\sigma})}2\tilde{\sigma}}{8\pi T}$$

$$\lambda_{2}^{(2\tilde{\sigma})} = -\frac{2\eta^{(2\tilde{\sigma})}2\tilde{\sigma}}{4\pi T} \int_{1}^{\infty} \frac{y^{2\tilde{\sigma}-2} - 1}{y(y^{2\tilde{\sigma}}-1)} dy$$

$$\lambda_{3}^{(2\tilde{\sigma})} = 0$$

Jet quenching parameter

Observed medium-induced suppression of back-to-back high-pT jets, at RHIC. Due to radiative energy loss of relativistic partons moving through the plasma.

Energy loss characterized by jet quenching parameter

$$\hat{q} = \frac{(\text{energy transf. from the plasma})^2}{\text{mean free path}}$$

Non perturbative definition in terms of a certain light-like Wilson loop. Evaluated in dual gravity setup [Liu, Rajagopal, Wiedemann 06]

$$\hat{q} = \frac{\pi^3 \sqrt{\lambda_h} \Gamma(\frac{3}{4})}{\sqrt{Vol(X_5)} \Gamma(\frac{5}{4})} T^3 \left[1 + \frac{1}{8} (2+\pi) \epsilon_h + \gamma \epsilon_h^2 + O(\epsilon_h^3) \right]$$
$$\gamma \approx 0.5565$$

Flavors enhance the jet quenching. Opposite in pQCD [Muller-Nagle 2006]